

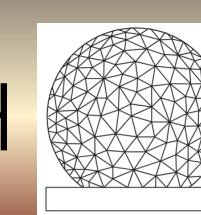
Challenges of Utilizing Mars' Resources to Warm Mars

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3 –step approach to terraforming

Near-term	Mid-term	Long-term
Today Mars is too cold and dry for Earth-like life to flourish. The first step is abiotic engineering to heat the planet.	A future, warmer Mars would be suitable for non-human life. A planetary ecosystem would begin producing oxygen.	In the long term, Mars would accumulate more atmosphere and have a stable, favorable climate.
Method: abiotic engineering GOAL: ☀ Temperature	Method: photosynthesis GOAL: 🌱 O ₂	Method: abiotic + biotic GOAL: 🌬 pressure, stabilize climate
THIS WORK		

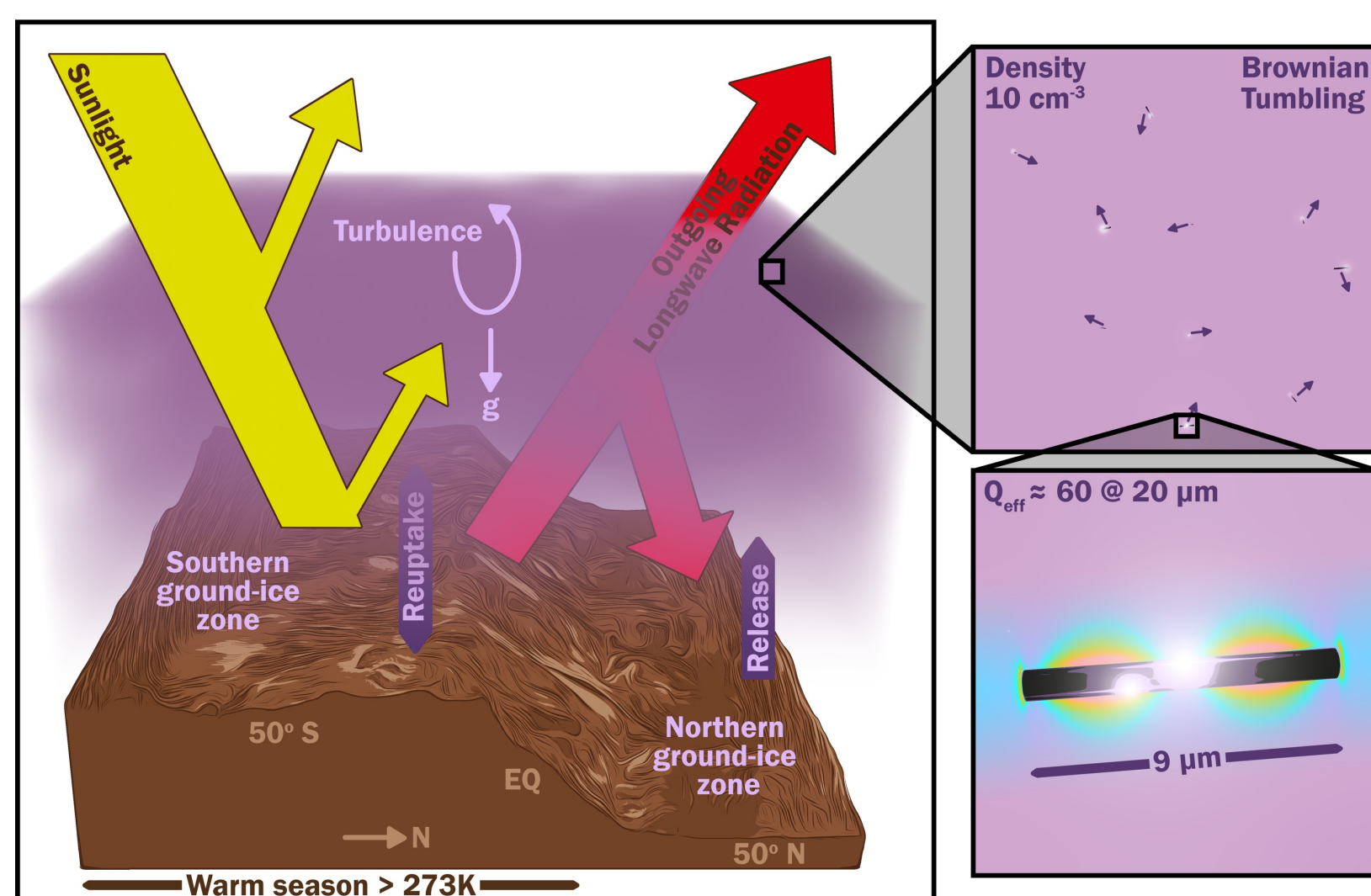
Warming Mars could be a step toward making it suitable for life, but would represent a major challenge for planetary science and engineering. Recent work suggests physically feasible methods [1,2,3], including engineered-aerosol warming [4,5]. However, before we can assess whether warming Mars is worthwhile, relative to the alternative of leaving Mars as a pristine wilderness, we must confront the practical requirements, cost, and possible risks. [6]

Approaches to warming Mars

The technological challenge of warming Mars can be approached from multiple angles. [7]

- Greenhouse gas, (e.g. CFC *chlorofluorocarbons*)
- Reflect energy from orbit mirrors, particles
- Diverting volatiles from asteroids
- Release CO₂ ice trapped in the caps (not sufficient by itself [8])
- Silica aerogel tiles [1]
- Engineered aerosols – x 5000 more efficient** per unit mass than gas [4]

Engineered aerosols provide favorable absorption and scattering of **visible light** and trap the **thermal radiation** from the surface.



Overview of the engineered-aerosol warming. Figure: A.M. Geller, Northwestern, CIERA + IT-RCDS. From [4].

Who we are



Astera Institute is a nonprofit research institute for public good based in Emeryville, in the San Francisco Bay area. Astera has a strong focus on open-science: we are working with partners at leading universities and institutions.



Edwin Kite



Alex Kling



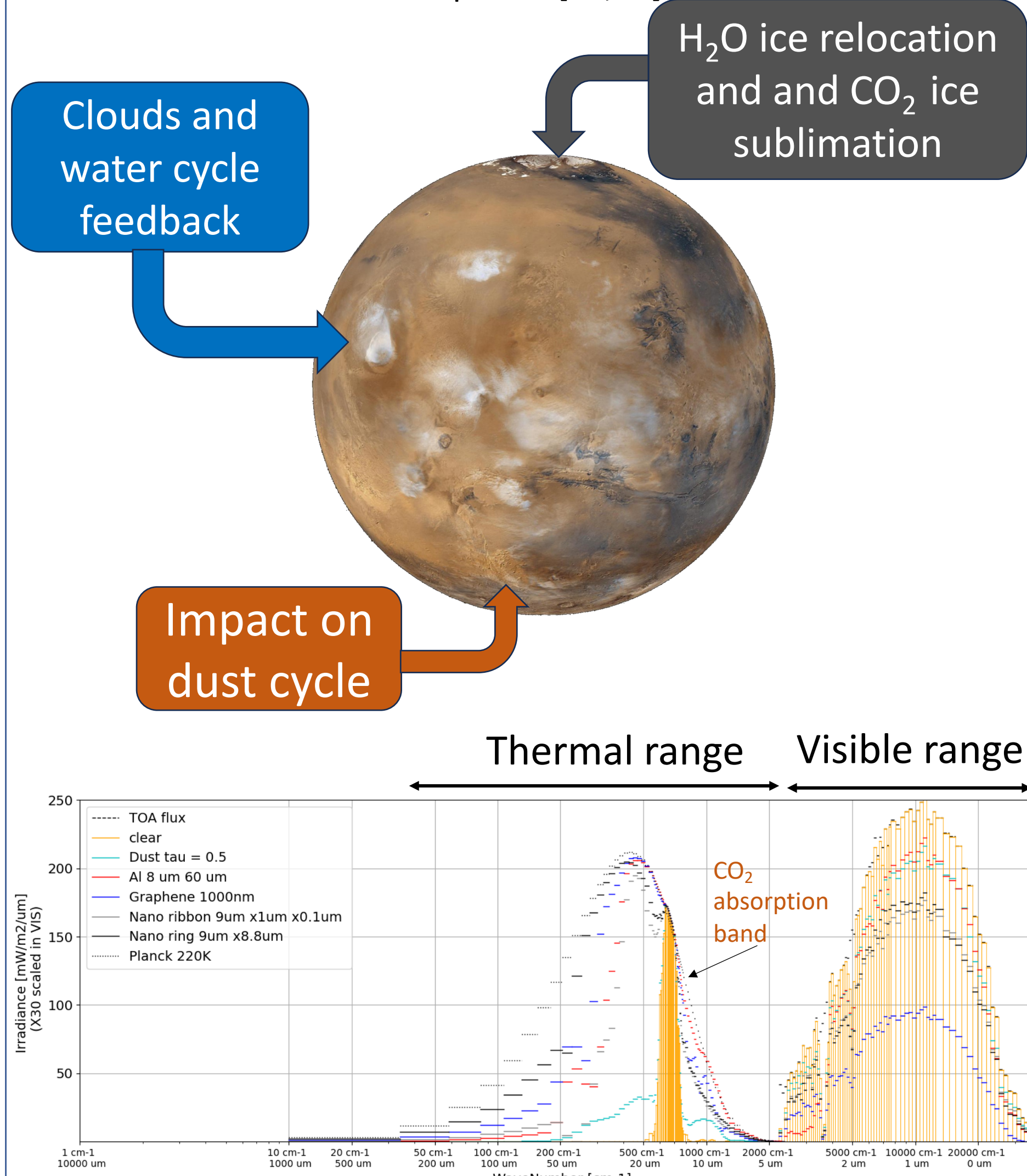
Ashwin Braude



Adrian Dumitrescu

Research

Understanding the **climate feedbacks** [9] due to the response of engineered aerosols is critical and best explored using 3D numerical models of the Martian atmosphere. [10,11]

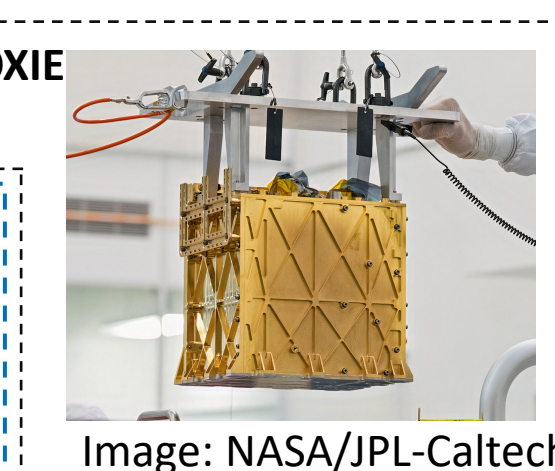
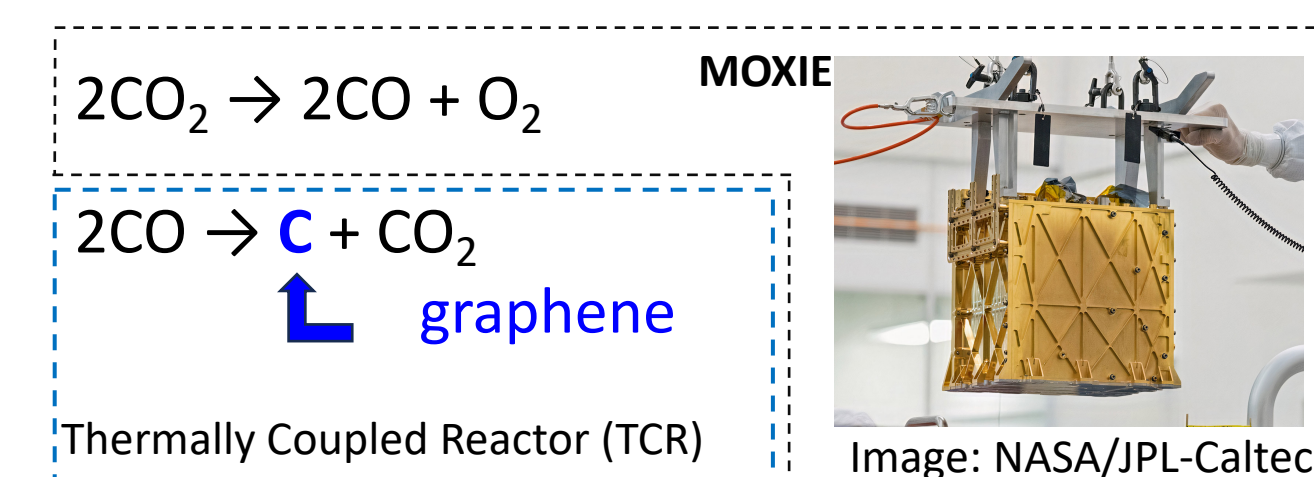


Our open-source particle screening tool calculates solar and thermal energy available at the surface for a clear atmosphere (**orange**), or with aerosols. github.com/alex-kling/TerraScreen

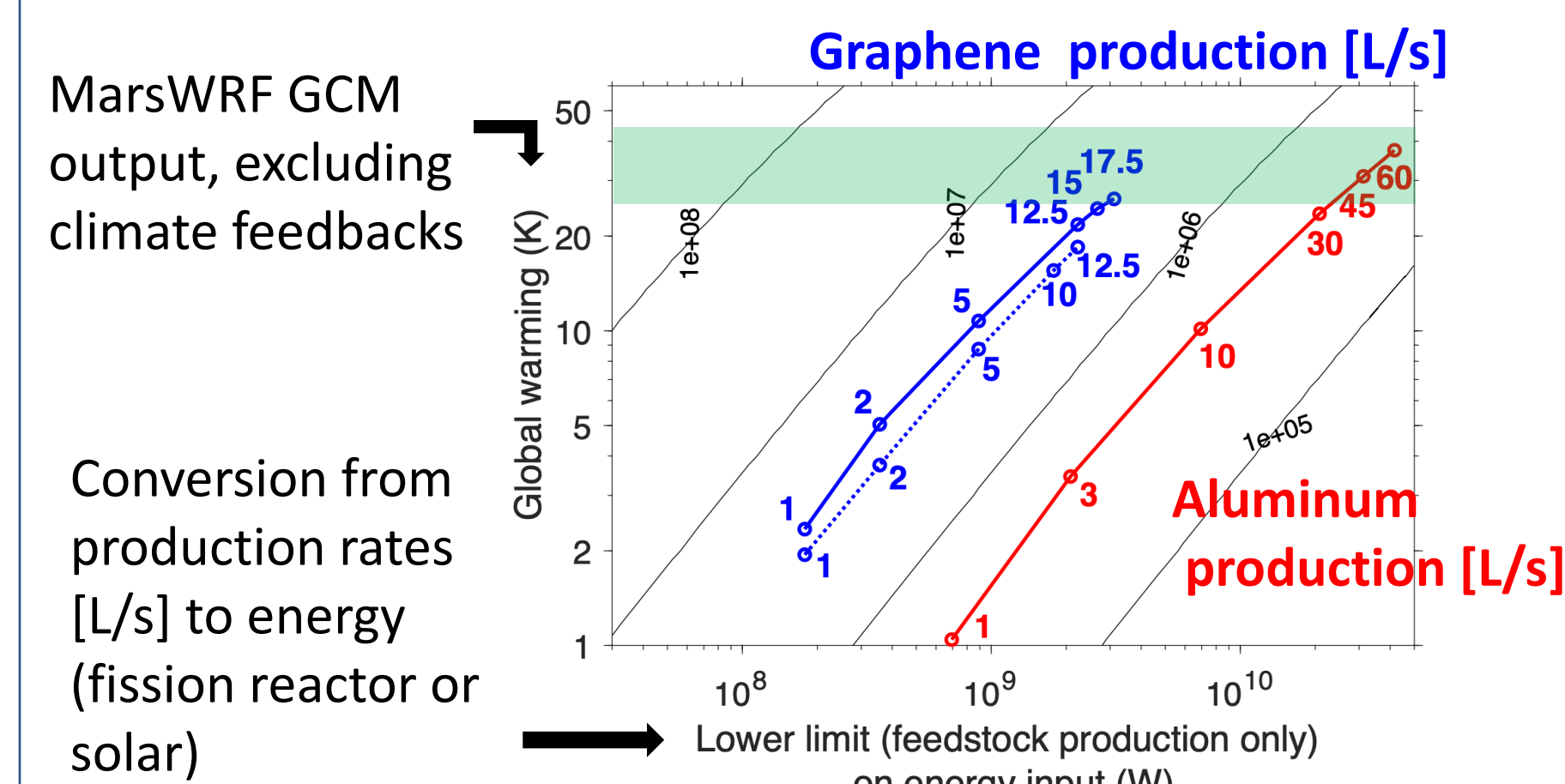
Open question: aerosol ISRU production

Example feedstock production for:

- Metallic nano-structures such as **Aluminum nano-rods**. Aluminum is **mined from the regolith**.
- Graphene** (i.e carbon) which is extracted from the CO₂ atmosphere through **solid oxide electrolysis (SOE)** [12,13,14,15]



Production of graphene through Solid Oxide Electrolysis (notional)

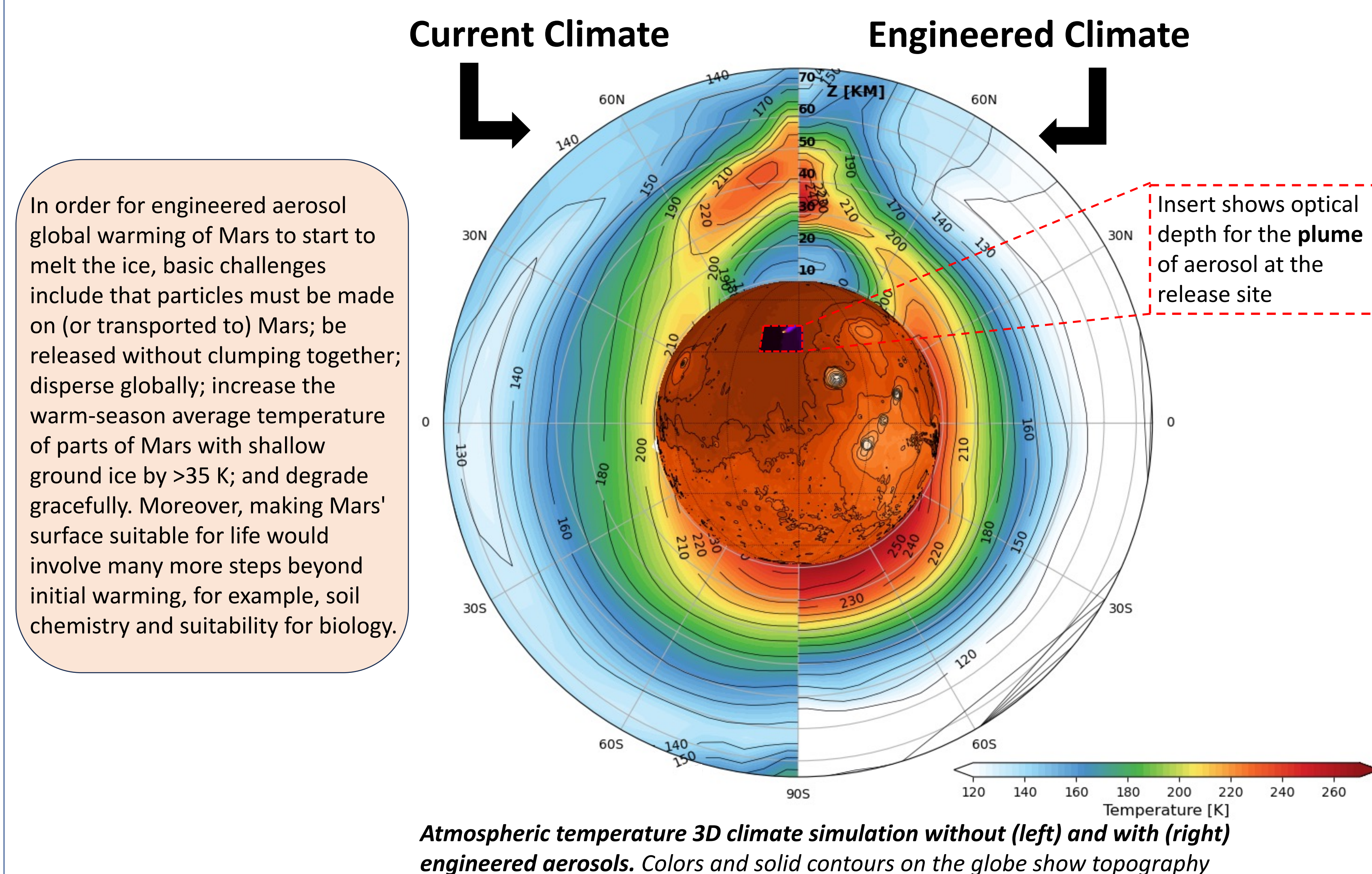


Conversion from production rates [L/s] to energy (fission reactor or solar)

Steady-state temperature response as a function of energy required to produce feedstock, for a mix of graphene disk of different diameters (blue), and for thin Al rods (red). Black contours show energy gain.

Initial results

- Global aerosols production rates of 10's L/s provide ~30K of warming within a few Earth years. [4,5]

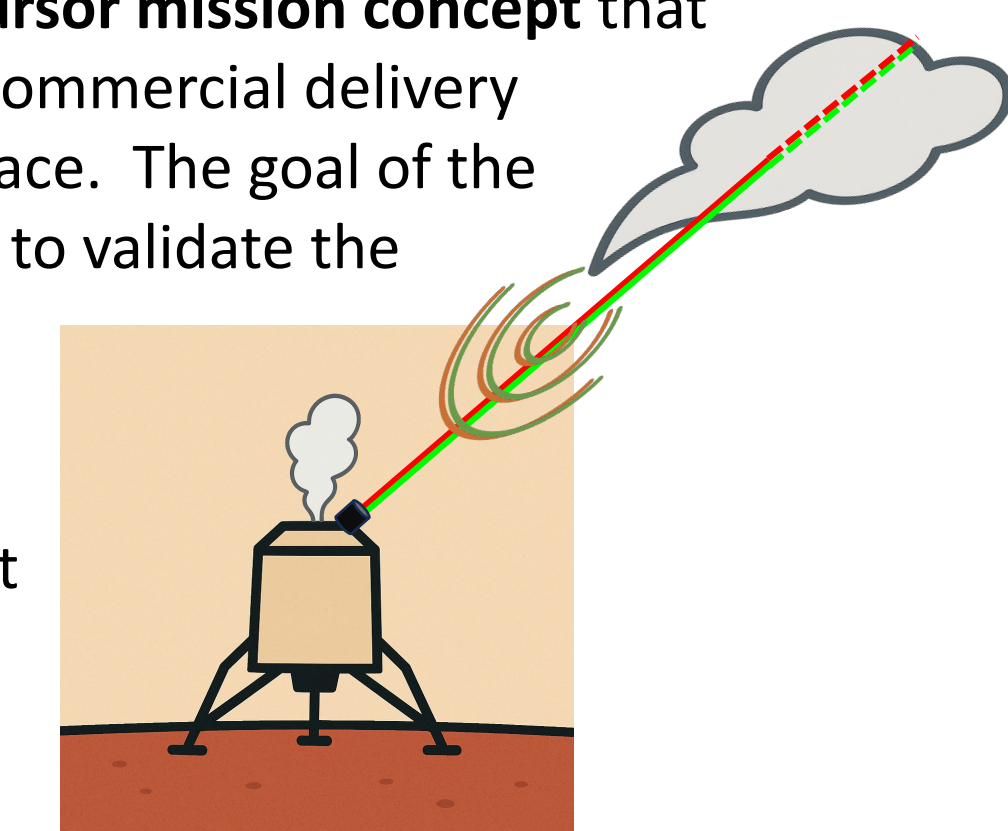


In order for engineered aerosol global warming of Mars to start to melt the ice, basic challenges include that particles must be made on (or transported to) Mars; be released without clumping together; disperse globally; increase the warm-season average temperature of parts of Mars with shallow ground ice by >35 K; and degrade gracefully. Moreover, making Mars' surface suitable for life would involve many more steps beyond initial warming, for example, soil chemistry and suitability for biology.

Atmospheric temperature 3D climate simulation without (left) and with (right) engineered aerosols. Colors and solid contours on the globe show topography

Upcoming work

- We are actively investigating **particle designs and trade-off** in terms of warming efficiency per unit mass and in-situ production at scale. Examples of other particles being assessed are:
 - Ø=9 µm nano ring (high efficiency, difficult production)
 - 9x1x0.1 µm nano ribbons (suitable for roll-to-roll production)
 - Coated particles (e.g. enhanced scattering properties)
 - Porous spheres (calculations from A. Raman's group, UCLA)
- **Particle manufacturing and validation.** We are in the process of manufacturing particles at a small scale (mg) to validate the simulated optical properties in a substrate using Fourier transform infrared spectroscopy (FTIR)
- **Particle dispersion tests.** We are setting-up a test chamber to characterize the deployment and clumping behavior of free floating nano-particles. The effort involves developing a backscatter differential absorption lidar (DIAL) to retrieve particles' position and size using remote sensing.
- We are investigating a **precursor mission concept** that would leverage upcoming commercial delivery services to the Martian surface. The goal of the precursor mission would be to validate the release of a small (<100 kg, non-climate -altering) amount of nano-particles to validate their deployment and dispersion in Martian conditions.



Precursor mission study to test the dispersal and tracking of particles using a backscatter lidar (notional)

Learn more

Greening the Solar System
Edwin Kite & Robin Wordsworth

Asterisk essay, 2025

Science Advances
Pilot study: Ansari et al.,
Science Advances, 2024

nature astronomy

Perspective paper: DeBenedictis et al. *Nature Astronomy*, 2025

Atmosphere Dynamics Impacts:
Richardson et al. 2025

References

- [1] Wordsworth R. et al. (2019) *Nat. Astronomy*, 3. [2] Handmer C. (2024) 10th Mars, #3025. [3] Wordsworth et al. (2025) *Astrobiology* 5. [4] Ansaris. et al. (2024) *Sci. Adv.*, 10. [5] Richardson et al. arXiv:2504.01455 (2025) [6] DeBenedictis E. et al. (2025) *Nat. Astronomy*, [7] Zubrin and McKay, (2005), 29th Joint Propulsion Conference and Exhibit . [8] Jakosky and Edwards, (2018) *Nat Astro.* 2, 634–639 [9] Madeleine, J., et al. (2014) *GRL*, 41. [10] Richardson M. I. et al. (2007) *JGR-Planets*, 112. [11] Toigo A. et al. (2012) *Icarus*, 221. [12] Hoffman J. A. et al. (2022) *Sci. Adv.*, 8 [13] Rapp D. & Hinterman E. (2023) *Space: Sci. &Tech.*, 3 [14] Hinterman, E., PhD thesis, MIT, 2022. [15] Grebenko A.K. et al. (2022) *Advanced Science* 9.